#### T2KK sensitivity as a function of off-axis angle

Fanny Dufour, Boston University

2<sup>nd</sup> T2KK workshop, July 13, 2006

## Outline

#### Motivation

#### Likelihood analysis:

- Analysis strategy
- Likelihood variables
- Efficiency results
- Future plans

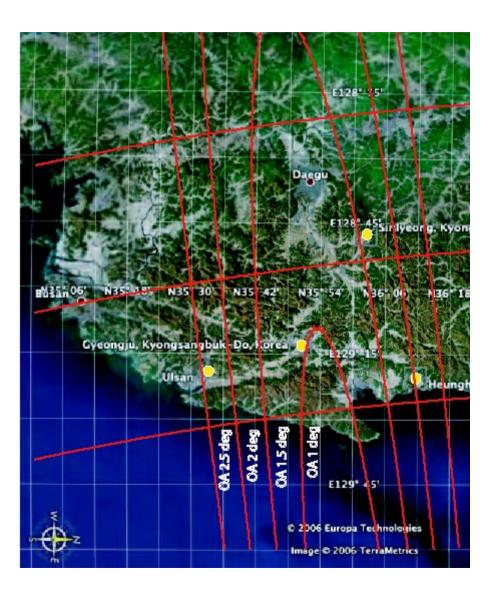
#### **Oscillation analysis**

- Introduction
- Spectrum (each off-axis angle)
- $\chi^2$  analysis
- Sensitivity curves
- Conclusions

## Overview

Study the sensitivity to CP violation and mass hierarchy as a function of the off-axis angle.

Axis considered: 1°Off-Axis (OA) 1.5°Off-Axis (OA) 2°Off-Axis (OA) 2.5°Off-Axis (OA)



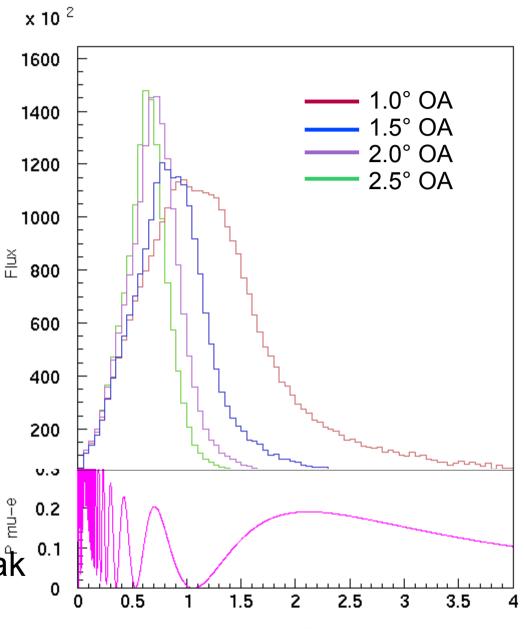
### Pros & cons

Small off-axis angle: (high energy tail)

1<sup>st</sup> appearance peak
 x more NC background

Big off-axis angle: (narrow peak)

Low background
 Low statistics at high E <sup>1</sup>/<sub>2</sub>
 Only 2<sup>nd</sup> appearance peak



Neutrino Energy

### Likelihood analysis strategy

Based on the T2K  $v_{a}$  appearance analysis

- Apply following precuts: FCFV, Evis <100 MeV Single ring e-like no decay electron
- In this study, I used the T2K Monte Carlo.
- Combine Super-K variables into a likelihood to discriminate electrons from  $\pi^0$ .

### 8 Variables

#### **Standard SK variables:**

Ring parameter PID parameter

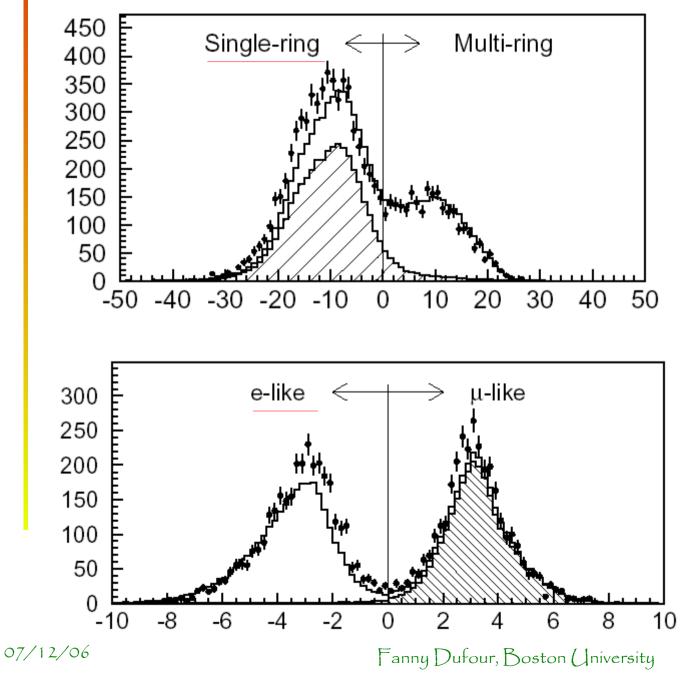
#### Special π° fitter variables: (POLfit, Pattern Of Light)

 $\pi^{\circ}$  mass  $\pi^{\circ}$  likelihood Energy fraction of 2<sup>nd</sup> ring

# New variables, defined for this analysis: Beam related variable:

Chi\_Xalong Chi\_cos(open)  $\cos\theta_{ve}$ 

## **Ring and PID Parameter:**



Those variables are not only precuts, (keep single-ring, e-like)

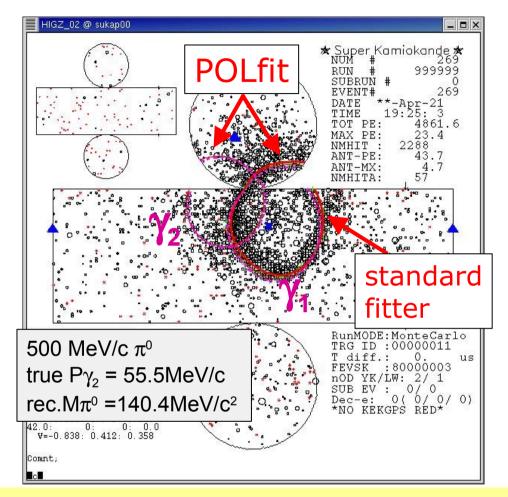
we also use the variables themselves in the likelihood.

MC
CCQE
SK data

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## POLfit

S.Nakayama's talk 1<sup>st</sup> T2KK Workshop November 2005



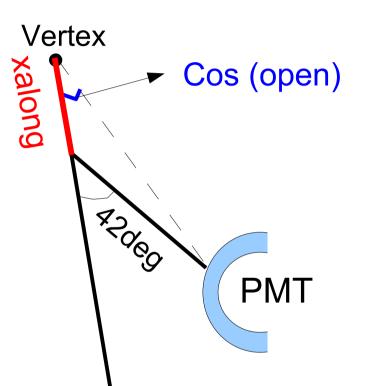
I use :  $\pi^{\circ}$  mass  $\pi^{\circ} \Delta Likelihood$ Energy fraction of 2<sup>nd</sup> ring

- Target: FCFV 1R-elike events
- ▲L≡Likelihood(2γ assump.) Likelihood(electron assump.)
- Try to reconstruct two  $\gamma$  rings
- Input: vertex, visible energy, and the  $1^{st} \gamma$  direction by the standard fitter
- Compare observed & expected (direct+scatter) charge
- Vary the 2<sup>nd</sup> γ direction and the energy fraction until the best match found

## Xalong & Cos(open)

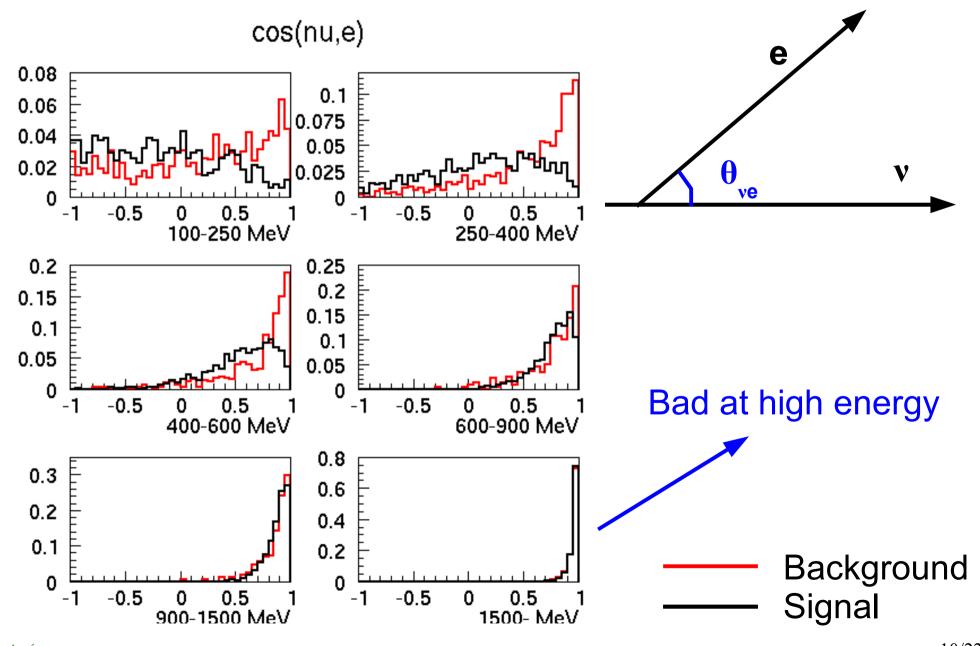
Xalong: Distance between vertex and emitting point of Cherenkov light.

Cos(open): Angle between vertex-pmt vector & direction of neutrino



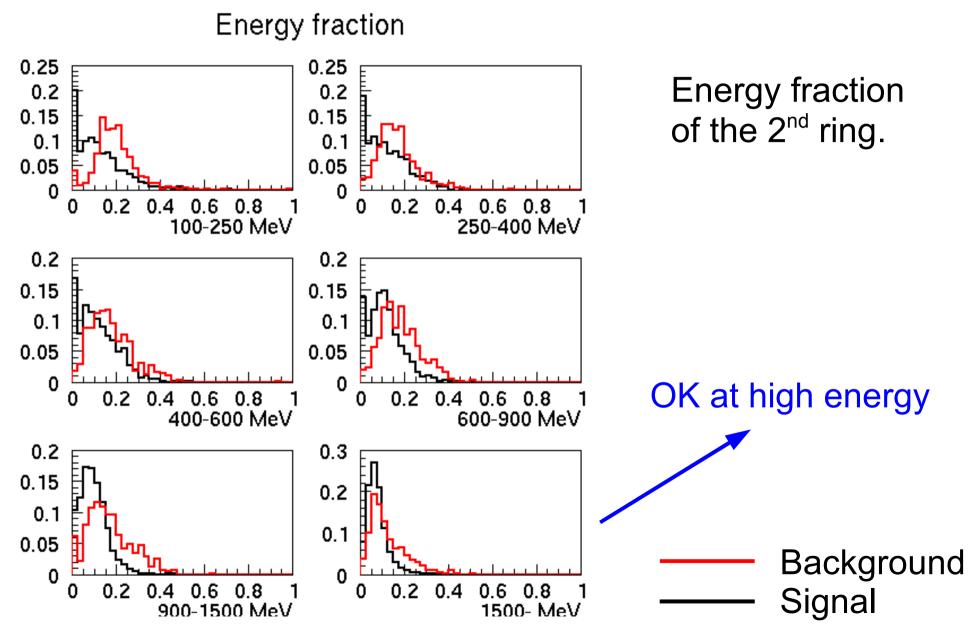
- I compute those values for each hit pmt, plot distributions.
- Using part of the MC I create templates of those distributions.
- For each event, I assign a  $\chi^2$  value comparing the event against the templates.
- The  $\chi^2$  value is added to the likelihood.

#### Example of distribution (1)



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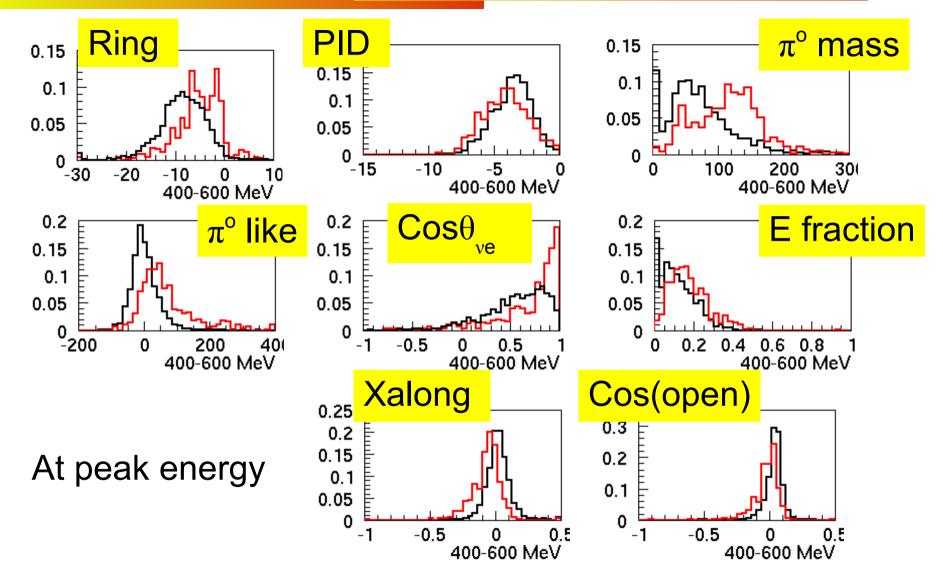
## Example of distribution (2)



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### **Overview of distributions:**



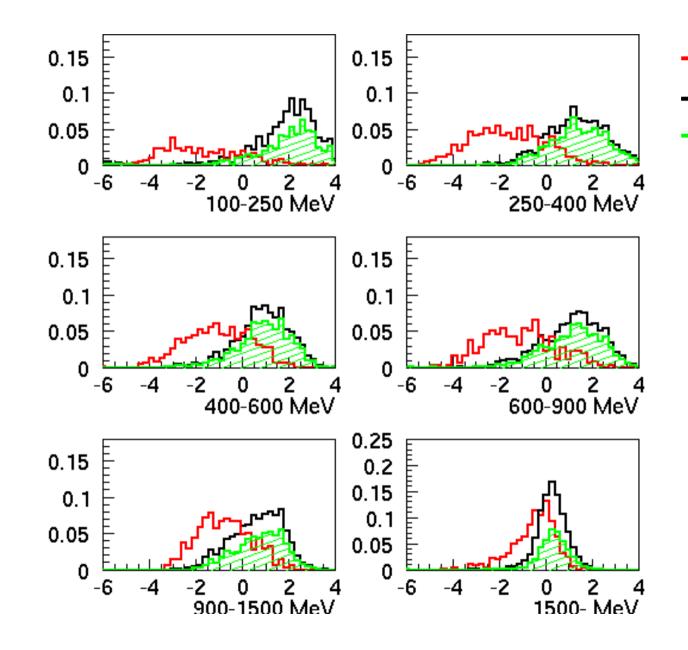
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Background

Signal

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## final likelihood



Background
Signal
v<sub>e</sub> CCQE

07/12/06

## **Final efficiency**

E <sub>rec</sub> (G	eV) 0-(	0.35	0.35-0.85	0.85-1.5	1.5-	NB:
ν <sub>μ</sub> CC	fcfv 2	86.9 70.2 3.6 1.4 0.2 14.6%	415.7 220.8 4.5 1.5 0.5 31.4%	370.4 146.3 5.3 1.9 0.6 32.0%	995.0 433.6 25.4 11.9 2.2 18.7%	arbitrary numbers
NC	fcfv 1ring e-like nodecay-e likelihood efficiency	5.1	229.6 66.2 57.2 53.1 10.9 20.5%	86.0 26.0 24.9 20.8 4.0 <b>19.5%</b>	83.6 41.1 39.6 32.6 11.1 <b>34.0%</b>	
V <sub>e</sub>	1ring e-like nodecay-e likelihood		36.7 21.6 21.3 18.9 15.4 81.8%	33.7 16.9 16.8 14.5 11.3 <b>78.3%</b>	73.3 37.4 37.2 30.8 22.1 71.7%	14

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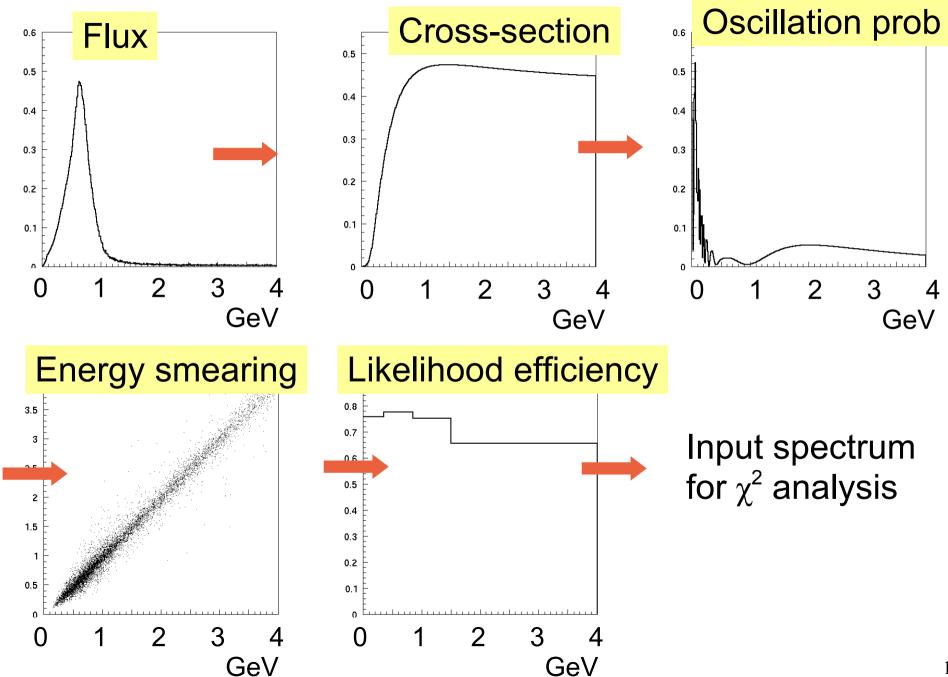
#### Likelihood future

There is room for improvement:

- Add new variables (Total pe charge/Evis, SK-II software variables, etc)
- Use different set of variables for different energies
- Extend analysis to higher energy bins
- Test Neural Network analysis
- Compare with atmospheric data

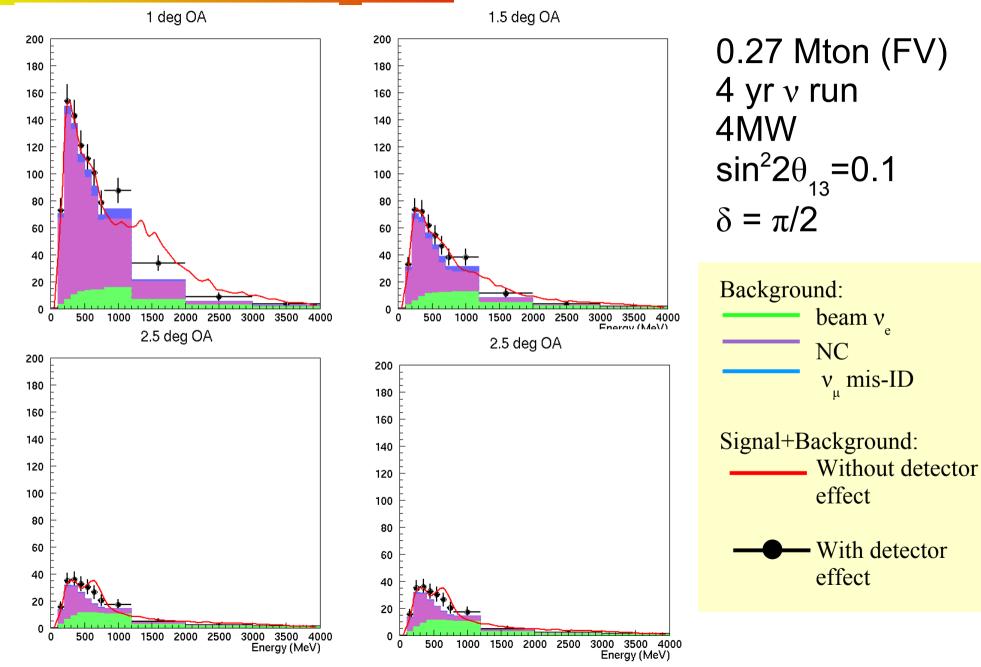
   → Check how well the variables are
   reproduced by MC.

## **Oscillation analysis**



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### **Spectrum for each OA**



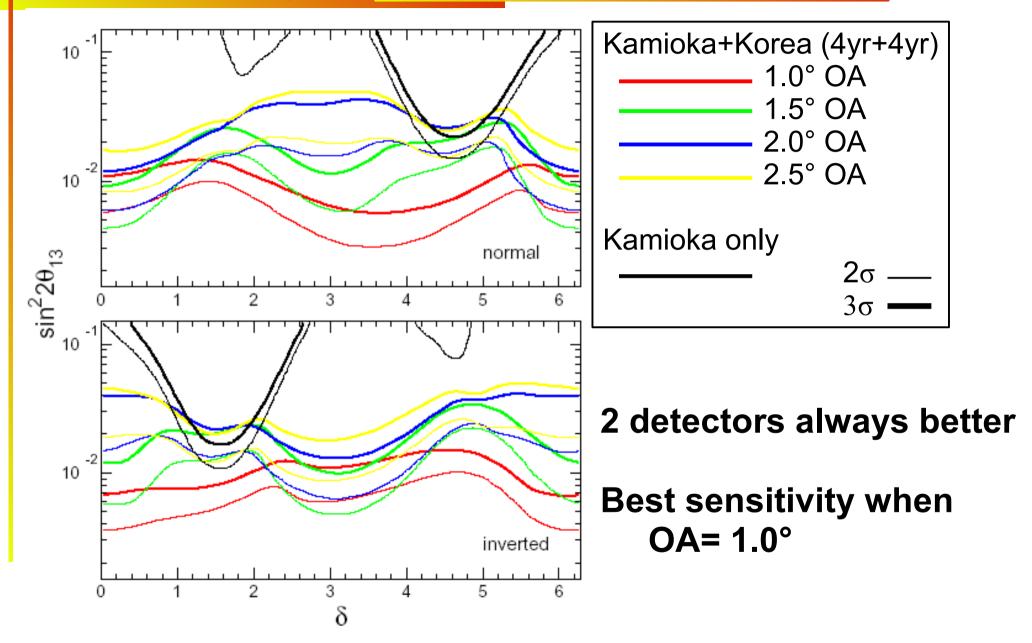
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## $\chi^2$ Definition

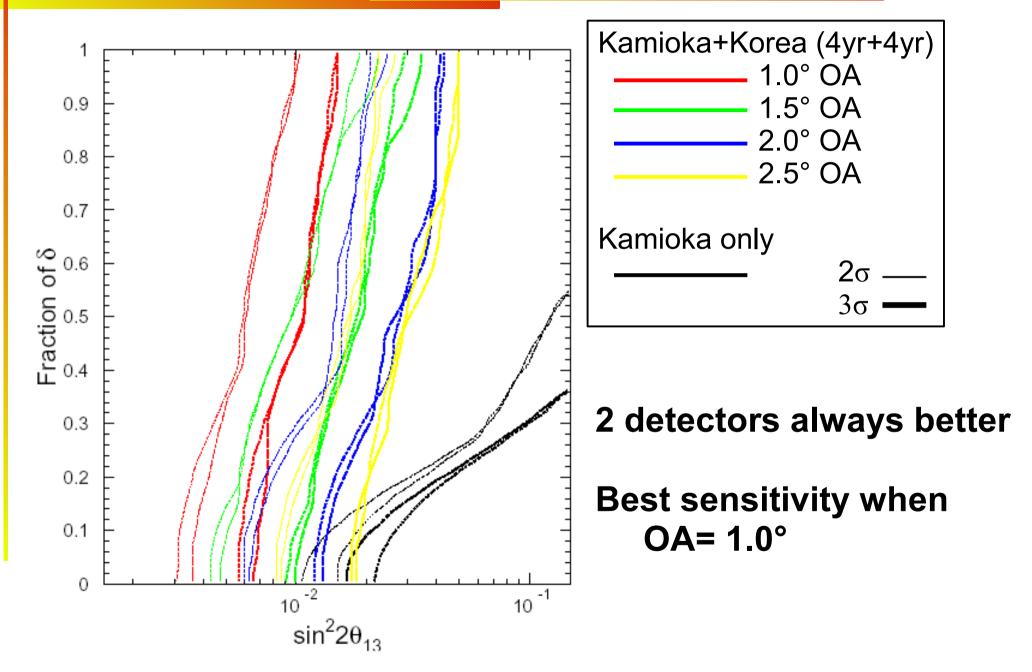
The oscillation analysis was done for: 4MW beam

k=1,4 0.27Mton in Korea 4 years running of neutrino 4 years running of antineutrino 6 years running of antineutrino With the following energy bins (MeV): i=1,7 400-500, 500-600, 600-700, 700-800, 800-1200,1200-2000, 2000-3000  $\chi^2 = \sum_{i=1}^{4} \left( \sum_{i=1}^{7} \frac{\left( N(e)_i^{\text{obs}} - N(e)_i^{\text{exp}} \right)^2}{\sigma_i^2} \right) + \sum_{i=1}^{3} \left( \frac{\epsilon_j}{\tilde{\sigma}_i} \right)^2$  $N(e)_i^{\text{exp}} = N_i^{\text{BG}} \cdot \left(1 + \sum_{j=1}^2 f_j^i \cdot \epsilon_j\right) + N_i^{\text{signal}} \cdot \left(1 + f_3^i \cdot \epsilon_3\right) \quad .$ hep-ph 0604026 eq 3) and 4)

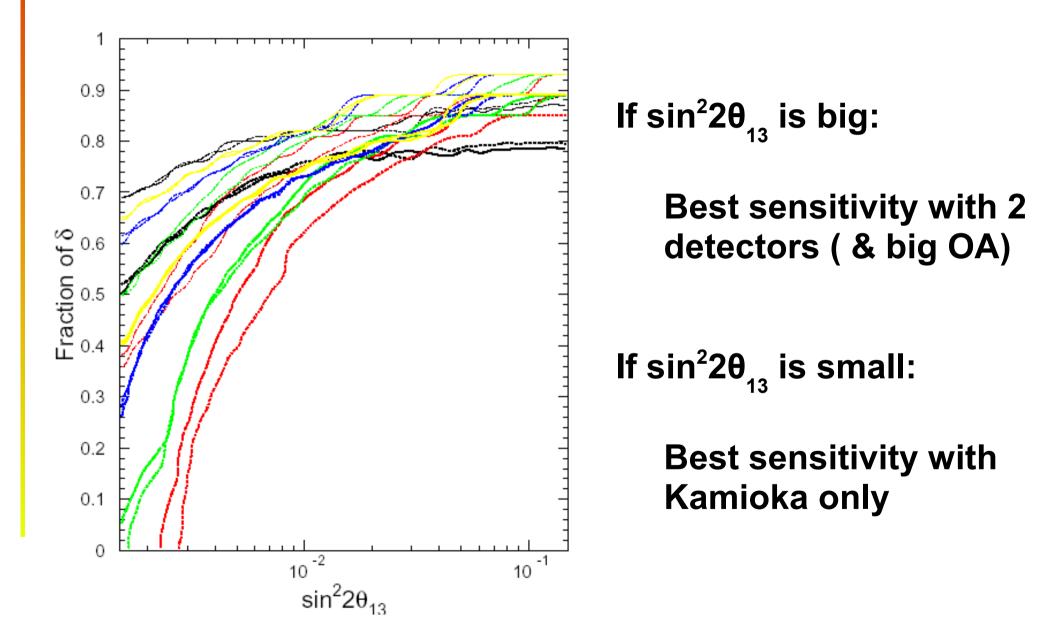
#### Sensitivity mass hierarchy



#### Sensitivity mass hierarchy



#### **Sensitivity CP violation**



#### **Conclusions**

Likelihood analysis developed for v appearance:

ε = 82% / BG = 21% → ε = 72% / BG = 34%

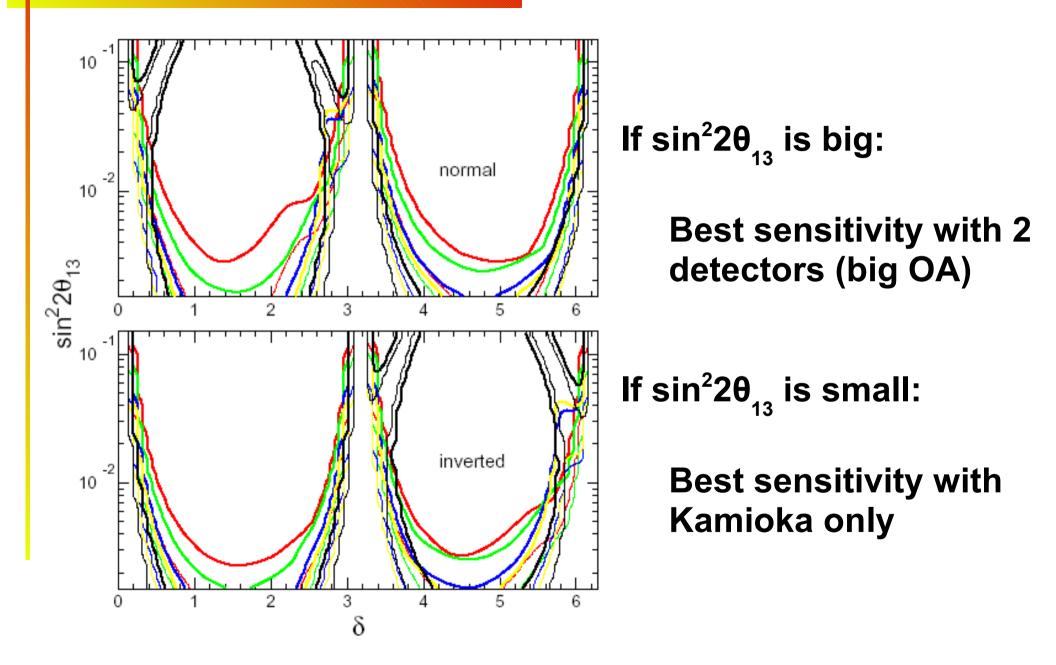
#### **Oscillation analysis conclusions:**

For CP violation study: Best set up is Kamioka only (for small  $\sin^2 2\theta_{13}$ ) or OA big (= 2.5°) if 2 detectors (for big  $\sin^2 2\theta_{13}$ )  $2^{nd}$  osc maximum  $\implies$  bigger CP effect

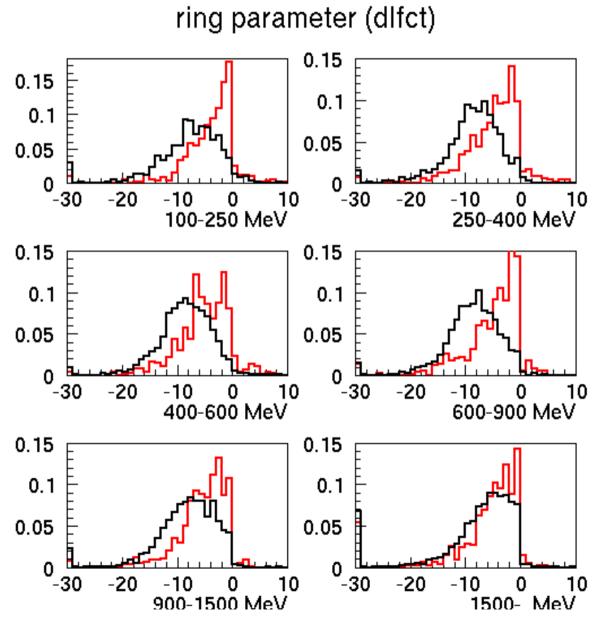
Future plan: Extend analysis to higher energies (especially for 1° OA)

#### **Backups:**

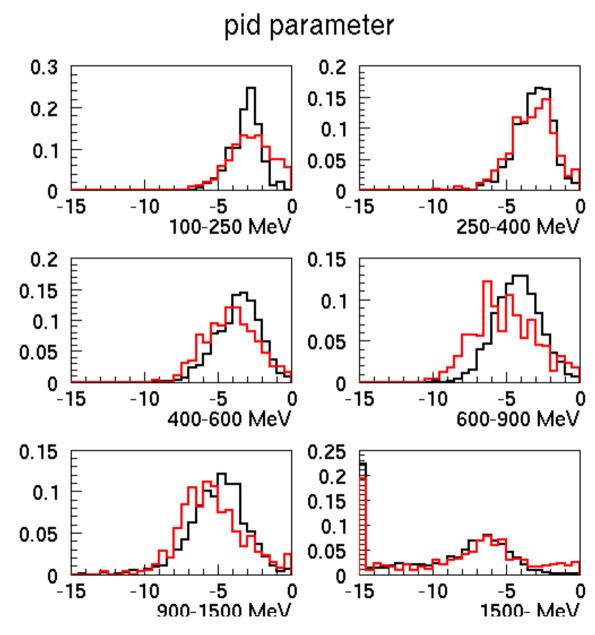
## **Sensitivity CP violation**



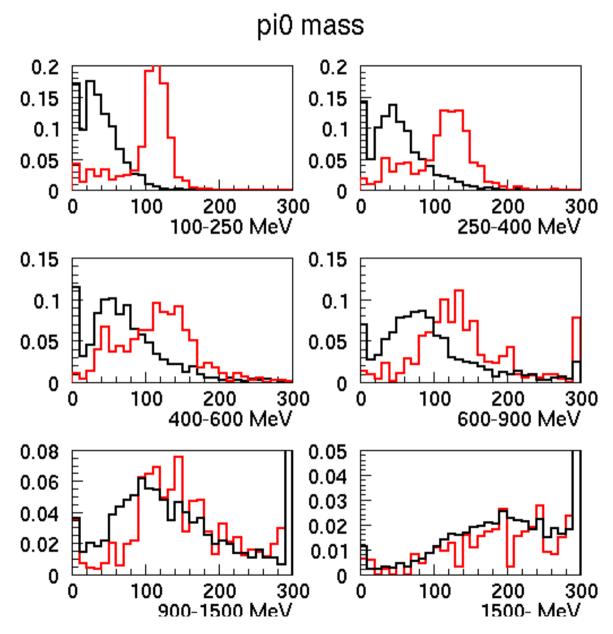
## ring param



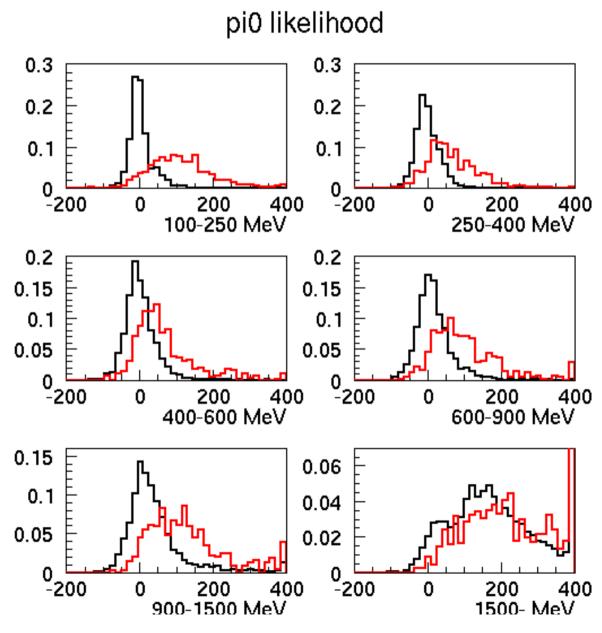
## pid



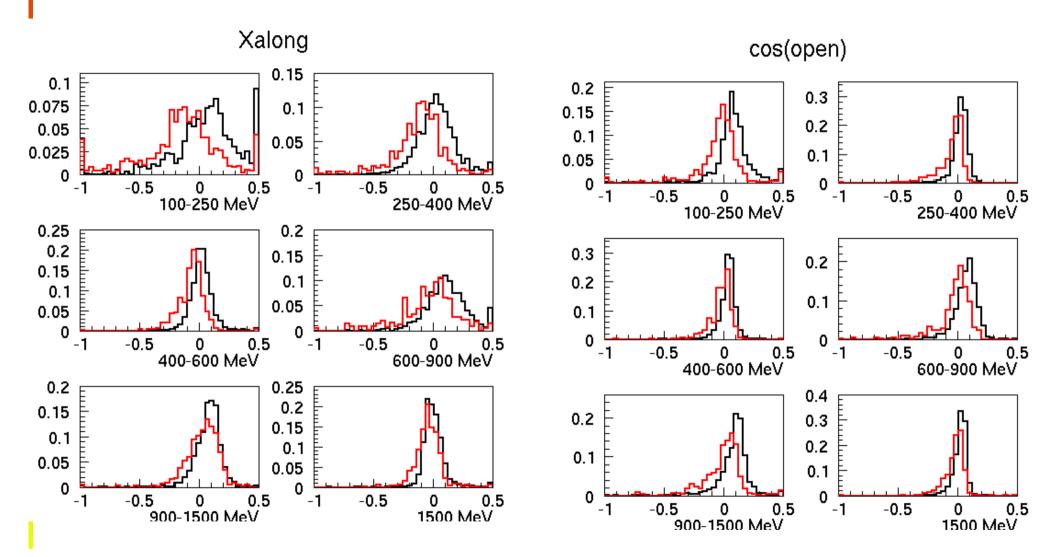
## pi0mass



## pi0like



## xalong cosopen (distribution)



#### usefulness of variables

add/remove variables eff tables