#### T2KK sensitivity as a function of off-axis angle

#### Updated version, August 23 2006

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## Outline

#### Motivation

#### Likelihood analysis:

- Analysis strategy
- Likelihood variables
- Efficiency results
- Future plans

#### **Oscillation analysis**

- Introduction
- Spectrum (each off-axis angle)
- $\chi^2$  analysis
- Sensitivity curves
- Conclusions

## **Overview**

Study the sensitivity to CP violation and mass hierarchy as a function of the off-axis angle.

Axis considered: 1°Off-Axis (OA) 1.5°Off-Axis (OA) 2°Off-Axis (OA) 2.5°Off-Axis (OA)



### Pros & cons

Small off-axis angle: (high energy tail)

1<sup>st</sup> appearance peak
 more NC background

Big off-axis angle: (narrow peak)

Low background
 Low statistics at high E <sup>1</sup>/<sub>E</sub>
 Only 2<sup>nd</sup> appearance peak



Neutrino Energy

#### Likelihood analysis strategy

Based on the T2K  $v_{a}$  appearance analysis

- Apply following precuts: FCFV, Evis >100 MeV Single ring e-like no decay electron
- In this study, I used the T2K Monte Carlo.
- Combine Super-K variables into a likelihood to discriminate electrons from  $\pi^0$ .

#### 8 Variables

#### **Standard SK variables:**

Ring parameter PID parameter

Chi Xalong

Chi cos(open)

#### Special π<sup>°</sup> fitter variables: (POLfit, Pattern Of Light)

 $\pi^{\circ}$  mass  $\pi^{\circ}$  likelihood Energy fraction of 2<sup>nd</sup> ring

#### New variables, defined for this analysis: Beam related variable:

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Cosθ

## **Ring and PID Parameter:**



Those variables are not only precuts, (keep single-ring, e-like)

we also use the variables themselves in the likelihood.

MC
CCQE
SK data

7/22

## POLfit



I use :  $\pi^{\circ}$  mass  $\pi^{\circ} \Delta Likelihood$ Energy fraction of 2<sup>nd</sup> ring

- Target: FCFV 1R-elike events
- ▲L≡Likelihood(2γ assump.) –
   Likelihood(electron assump.)
- Try to reconstruct two  $\gamma$  rings
- Input: vertex, visible energy, and the 1<sup>st</sup> γ direction by the standard fitter
- Compare observed & expected (direct+scatter) charge
- Vary the 2<sup>nd</sup> γ direction and the energy fraction until the best match found

# Xalong & Cos(open)

Xalong: Distance between vertex and emitting point of Cherenkov light.

Cos(open): Angle between vertex-pmt vector & direction of particle



- I compute those values for each hit pmt, plot distributions.
- Using part of the MC I create templates of those distributions.
- For each event, I assign a  $\chi^2$  value comparing the event against the templates.
- The  $\chi^2$  value is added to the likelihood.

#### Example of distribution (1)



### Example of distribution (2)



#### **Overview of distributions:**



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Background

Signal

### final likelihood



Background
Signal
v<sub>e</sub> CCQE

### **Final efficiency**

E <sub>rec</sub> (Ge	eV) 0-0	).35 (	).35-0.85	0.85-1.5	1.5-
ν <sub>μ</sub> CC	fcfv 2	86.9	415.7	370.4	995.0
	1ring 1	70.2	220.8	146.3	433.6
	e-like	3.6	4.5	5.3	25.4
	nodecay-e	1.4	1.5	1.9	11.9
	likelihood	0.2	0.5	0.6	2.2
	efficiency	14.6%	31.4%	32.0%	18.7%
NC	fcfv	22.0	229.6	86.0	83.6
	1ring	89.0	66.2	26.0	41.1
	e-like	53.4	57.2	24.9	39.6
	nodecay-e	50.4	53.1	20.8	32.6
	likelihood	5.1	10.9	4.0	11.1
	efficiency	10.1%	20.5%	<b>19.5%</b>	<b>34.0%</b>
V <sub>e</sub>	fcfv 1	2.2	36.7	33.7	73.3
	1ring	5.7	21.6	16.9	37.4
	e-like	5.6	21.3	16.8	37.2
	nodecay-e	4.7	18.9	14.5	30.8
	likelihood	4.0	15.4	11.3	22.1
	efficiency	85.4%	81.8%	78.3%	71.7%

NB: arbitrary numbers

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#### Likelihood future

There is room for improvement:

- Add new variables (Total pe charge/Evis, SK-II software variables, etc)
- Use different set of variables for different energies
- Extend analysis to higher energy bins
- Test Neural Network analysis
- Compare with atmospheric data

   → Check how well the variables are
   reproduced by MC.

## **Oscillation analysis**



#### **Spectrum for each OA**



## $\chi^2$ Definition

The oscillation analysis was done for: 4MW beam

k=1,4 0.27Mton in Korea 0.54Mton in Kamioka 0.27Mton in Korea *When Kamioka* 4 years running of neutrino *only* 4 years running of antineutrino With the following energy bins (MeV): i=1,7 400-500, 500-600, 600-700, 700-800, 800-1200,1200-2000, 2000-3000  $\chi^2 = \sum_{i=1}^{4} \left( \sum_{i=1}^{7} \frac{\left( N(e)_i^{\text{obs}} - N(e)_i^{\text{exp}} \right)^2}{\sigma_i^2} \right) + \sum_{i=1}^{3} \left( \frac{\epsilon_j}{\tilde{\sigma}_i} \right)^2$  $N(e)_i^{\text{exp}} = N_i^{\text{BG}} \cdot \left(1 + \sum_{j=1}^2 f_j^i \cdot \epsilon_j\right) + N_i^{\text{signal}} \cdot \left(1 + f_3^i \cdot \epsilon_3\right) \quad .$ hep-ph 0604026 eq 3) and 4)

#### One bug was fixed...

So the next set of slide always have and old and a new slide.

The bug was: In the case of the 2 detector setup (Kamioka+Korea) I was assigning the background according to the off-axis angle for both Kamioka and Korea, which is wrong.

I should have kept the background setup to 2.50A for Kamioka and change only the background for Korea.

Sorry for the confusion.

### Sensitivity mass hierarchy (old)



### Sensitivity mass hierarchy (new)



### Sensitivity mass hierarchy (old)



## Sensitivity mass hierarchy (new)



### Sensitivity CP violation (old)



## Sensitivity CP violation (new)



## **Conclusions (old)**

Likelihood analysis developed for v appearance:

ε = 82% / BG = 21% → ε = 72% / BG = 34%

#### **Oscillation analysis conclusions:**

For mass hierarchy: Best set up is when OA is small (= 1.0°) 1<sup>st</sup> osc maximum atter effect

For CP violation study: Best set up is Kamioka only (for small  $\sin^2 2\theta_{13}$ ) or OA big (= 2.5°) if 2 detectors (for big  $\sin^2 2\theta_{13}$ )  $2^{nd}$  osc maximum  $\longrightarrow$  bigger CP effect

Future plan: Extend analysis to higher energies (especially for 1° OA)

## **Conclusions (new)**

Likelihood analysis developed for v appearance:

ε = 82% / BG = 21% → ε = 72% / BG = 34%

#### **Oscillation analysis conclusions:**

For mass hierarchy: Best set up is when OA is small (= 1.0°) 1<sup>st</sup> osc maximum \_\_\_\_\_ matter effect

#### For CP violation study:

Not many difference for different OA angle unless  $\theta_{13}$  is very small and in that case 1.0° OA is the best of 2 detector setup, but Kamioka only would even be better.

# Future plan: Extend analysis to higher energies (especially for 1° OA)

#### **Backups:**

### **Sensitivity CP violation**



# ring param



## pid



## pi0mass



## pi0like



### xalong cosopen (distribution)



#### usefulness of variables

add/remove variables eff tables